Levinson's theorem for non-local interactions in two dimensions

This article has been downloaded from IOPscience. Please scroll down to see the full text article.
1998 J. Phys. A: Math. Gen. 317501
(http://iopscience.iop.org/0305-4470/31/37/010)
View the table of contents for this issue, or go to the journal homepage for more

Download details:
IP Address: 171.66.16.102
The article was downloaded on 02/06/2010 at 07:12

Please note that terms and conditions apply.

# Levinson's theorem for non-local interactions in two dimensions 

Shi-Hai Dong $\dagger \|$, Xi-Wen Hou $\ddagger \ddagger$ and Zhong-Qi Ma $\dagger \S$<br>$\dagger$ Institute of High Energy Physics, PO Box 918(4), Beijing 100039, People's Republic of China $\ddagger$ Department of Physics, University of Three Gorges, Yichang 443000, People’s Republic of China<br>§ China Center for Advanced Science and Technology (World Laboratory), PO Box 8730, Beijing 100080, People's Republic of China

Received 19 March 1998, in final form 29 June 1998


#### Abstract

In the light of the Sturm-Liouville theorem, the Levinson theorem for the Schrödinger equation with both local and non-local cylindrically symmetric potentials is studied. It is proved that the two-dimensional Levinson theorem holds for the case with both local and non-local cylindrically symmetric cut-off potentials, which is not necessarily separable. In addition, the problems related to the positive-energy bound states and the physically redundant state are also discussed.


## 1. Introduction

The Levinson theorem [1], an important theorem in scattering theory, established the relation between the total number of bound states and the phase shifts at zero momentum. The Levinson theorem has been proved by several authors with different methods, and generalized to different fields [2-9], including the cases with the non-local interactions [4, 5]. Generally speaking, three main methods are used to prove the Levinson theorem. One [1] is based on an elaborate analysis of the Jost function. This method requires a well-behaved potential. The second, based on the Green-function method [6], expounds that the total number of the physical states, which is infinite, is proved to be independent of the potential and the number of the bound states is the difference between the infinite numbers of the scattering states without and with the potential. The third method proves the Levinson theorem by the Sturm-Liouville theorem [7, 8]. This simple, intuitive method can easily be generalized. Some obstacles and ambiguities, which may occur with the other two methods, disappear in the third method. We have succeeded in dealing with the non-relativistic and relativistic problems in two dimensions in this way [10, 11].

The reasons why we write this paper are that, on the one hand, the Levinson theorem in two dimensions has been studied by experiment [12] as well as theoretically [13, 10, 11] because of the wide interest in lower-dimensional field theories and other modern physics [14-20], and on the other, the Levinson theorem for non-local interactions in two dimensions has never appeared in the literature.

This paper is organized as follows. In section 2, we establish the Sturm-Liouville theorem for non-local interactions in two dimensions. The Levinson theorem for this case

[^0]will be set up in section 3. Some problems related to the positive-energy bound states and the physically redundant state will be studied in sections 4 and 5 .

## 2. The Sturm-Liouville theorem

Throughout this paper $\hbar=1$ and the mass $\mu=\frac{1}{2}$ are employed for simplicity. Consider the Schrödinger equation with a local potential $V(r)$ and a non-local potential $U\left(r, r^{\prime}\right)$, where both potentials are cylindrically symmetric:

$$
\begin{align*}
H \psi(r, \varphi)= & -\left(\frac{1}{r} \frac{\partial}{\partial r} r \frac{\partial}{\partial r}+\frac{1}{r^{2}} \frac{\partial^{2}}{\partial \varphi^{2}}\right) \psi(r, \varphi)+V(r) \psi(r, \varphi) \\
& +\int U\left(r, r^{\prime}\right) \delta\left(\varphi^{\prime}-\varphi\right) \psi\left(r^{\prime}, \varphi^{\prime}\right) r^{\prime} \mathrm{d} r^{\prime} \mathrm{d} \varphi \\
= & E \psi(r, \varphi) . \tag{1}
\end{align*}
$$

The mesonic theory of nuclear forces indicates that the interaction between two nucleons is local at great distances but becomes non-local if the two nucleons come close. To simplify the expression, we assume, following Martin [4], that the non-local potential $U\left(r, r^{\prime}\right)$ is real, continuous, symmetric, and vanishing at large distances [21]:

$$
U\left(r, r^{\prime}\right)= \begin{cases}U\left(r^{\prime}, r\right) &  \tag{2}\\ U\left(r, r^{\prime}\right)=\mathrm{O}\left(r^{-1}\right) & \text { at } r \sim 0 \\ U\left(r, r^{\prime}\right)=0 & \text { when } r \geqslant r_{0}\end{cases}
$$

As usual, the local potential $V(r)$ is real and continuous. We assume that $V(r)$ satisfies

$$
V(r)= \begin{cases}\mathrm{O}\left(r^{-1}\right) & \text { at } r \sim 0  \tag{3}\\ V(r)=0 & \text { when } r \geqslant r_{0}\end{cases}
$$

The first condition is necessary for the wavefunction to be well-behaved at the origin [1], and the potential with the second condition is called a cut-off one, namely, it vanishingly small beyond a sufficiently large radius $r_{0}$. It was proved that the tail of the local potential at infinity will not change the essence of the proof [10] if it decays faster than $r^{-3}$ at infinity. Under this assumption, the range of the in (1) is, in fact, from 0 to $r_{0}$, and the equation in the region $\left[r_{0}, \infty\right)$ becomes that for the free particle.

We introduce a parameter $\lambda$ for the potentials

$$
\begin{equation*}
V(r, \lambda)=\lambda V(r) \quad U\left(r, r^{\prime}, \lambda\right)=\lambda U\left(r, r^{\prime}\right) \tag{4}
\end{equation*}
$$

As $\lambda$ increases from zero to one, the potentials $V(r, \lambda)$ and $U\left(r, r^{\prime}, \lambda\right)$ change from zero to the given potentials $V(r)$ and $U\left(r, r^{\prime}\right)$, respectively.

Owing to the symmetry of the potentials, letting

$$
\begin{equation*}
\psi(r, \phi, \lambda)=r^{-1 / 2} R_{E m}(r, \lambda) \mathrm{e}^{ \pm \mathrm{i} m \phi} \quad m=0,1,2, \ldots \tag{5}
\end{equation*}
$$

we obtain the radial equation

$$
\begin{align*}
\frac{\partial^{2}}{\partial r^{2}} R_{E m}(r, \lambda) & +\left(E-V(r, \lambda)-\frac{m^{2}-1 / 4}{r^{2}}\right) R_{E m}(r, \lambda) \\
= & \sqrt{r} \int U\left(r, r^{\prime}, \lambda\right) R_{E m}\left(r^{\prime}, \lambda\right) \sqrt{r^{\prime}} \mathrm{d} r^{\prime} \tag{6}
\end{align*}
$$

where $\pm m$ and $E$ denote the angular momentum and energy, respectively. Since the radial function $R_{E m}(r, \lambda)$ is independent of the sign of the angular momentum, we only discuss the cases with non-negative $m$.

We are going to solve the radial equation (6) in two regions $\left[0, r_{0}\right]$ and $\left[r_{0}, \infty\right)$, and match two solutions at $r_{0}$. Since the Schrödinger equation is linear, the wavefunction can be multiplied by a constant factor. Removing the effect of the factor, we only need one matching condition at $r_{0}$ for the logarithmic derivative of the radial function:
$A_{m}(E, \lambda) \equiv\left\{\frac{1}{R_{E m}(r, \lambda)} \frac{\partial R_{E m}(r, \lambda)}{\partial r}\right\}_{r=r_{0}-}=\left\{\frac{1}{R_{E m}(r, \lambda)} \frac{\partial R_{E m}(r, \lambda)}{\partial r}\right\}_{r=r_{0}+}$.
We now turn to the Sturm-Liouville theorem. We denote by $\bar{R}_{E m}(r, \lambda)$ a solution of (6) with the energy $\bar{E}$. Multiplying the equations for $R_{E m}(r, \lambda)$ and $\bar{R}_{E m}(r, \lambda)$ by $\bar{R}_{E m}(r, \lambda)$ and $R_{E m}(r, \lambda)$, respectively, and calculating their difference, we have

$$
\begin{align*}
\frac{\partial}{\partial r}\left(R_{E m}(r, \lambda)\right. & \left.\frac{\partial \bar{R}_{E m}(r, \lambda)}{\partial r}-\bar{R}_{E m}(r, \lambda) \frac{\partial R_{E m}(r, \lambda)}{\partial r}\right)+(\bar{E}-E) R_{E m}(r, \lambda) \bar{R}_{E m}(r, \lambda) \\
= & \sqrt{r} R_{E m}(r, \lambda) \int U\left(r, r^{\prime}, \lambda\right) \bar{R}_{E m}\left(r^{\prime}, \lambda\right) \sqrt{r^{\prime}} \mathrm{d} r^{\prime} \\
& -\sqrt{r} R_{E m}(r, \lambda) \int U\left(r, r^{\prime}, \lambda\right) R_{E m}\left(r^{\prime}, \lambda\right) \sqrt{r^{\prime}} \mathrm{d} r^{\prime} \tag{8}
\end{align*}
$$

According to the boundary condition, both $R_{E m}(r, \lambda)$ and $\bar{R}_{E m}(r, \lambda)$ go to zero when $r$ tends to zero. Integrating (8) over the variable $r$ in the region $\left[0, r_{0}\right]$ and noting the symmetric property of $U\left(r, r^{\prime}\right)$, we have

$$
\begin{gathered}
\frac{1}{\bar{E}-E}\left\{R_{E m}(r, \lambda) \frac{\partial \bar{R}_{E m}(r, \lambda)}{\partial r}-\bar{R}_{E m}(r, \lambda) \frac{\partial R_{E m}(r, \lambda)}{\partial r}\right\}_{r=r_{0}-} \\
=-\int_{0}^{r_{0}} R_{E m}\left(r^{\prime}, \lambda\right) \bar{R}_{E m}\left(r^{\prime}, \lambda\right) \mathrm{d} r^{\prime}
\end{gathered}
$$

Taking the limit, we obtain

$$
\begin{align*}
\frac{\partial A_{m}(E, \lambda)}{\partial E} & \equiv \frac{\partial}{\partial E}\left(\frac{1}{R_{E m}(r, \lambda)} \frac{\partial R_{E m}(r, \lambda)}{\partial r}\right)_{r=r_{0}-} \\
& =-R_{E m}\left(r_{0}, \lambda\right)^{-2} \int_{0}^{r_{0}} R_{E m}\left(r^{\prime}, \lambda\right)^{2} \mathrm{~d} r^{\prime}<0 \tag{9}
\end{align*}
$$

Similarly, from the boundary condition that when $E \leqslant 0$ the radial function $R_{E m}(r, \lambda)$ tends to zero at infinity, we have

$$
\begin{equation*}
\frac{\partial}{\partial E}\left(\frac{1}{R_{E m}(r, \lambda)} \frac{\partial R_{E m}(r, \lambda)}{\partial r}\right)_{r=r_{0}+}=R_{E m}\left(r_{0}, \lambda\right)^{-2} \int_{r_{0}}^{\infty} R_{E m}\left(r^{\prime}, \lambda\right)^{2} \mathrm{~d} r^{\prime}>0 \tag{10}
\end{equation*}
$$

Therefore, when $E \leqslant 0$, both sides of the matching condition (7) are monotonic with respect to the energy $E$. As energy increases, the logarithmic derivative of the radial function at $r_{0}-$ decreases monotonically, but that at $r_{0}+$ increases monotonically. This is an expression of the Sturm-Liouville theorem [22].

## 3. The Levinson theorem

The establishment of the Levinson theorem for the case with both local and non-local cylindrically symmetric potentials is similar to that for the case with only a local potential.

In solving the radial equation (6) in the region $\left[0, r_{0}\right]$, only one solution is convergent at the origin. Thus, for the given potentials the logarithmic derivative $A_{m}(E, \lambda)$ is determined in principle. For example, for free particle $(\lambda=0)$ we have
$R_{E m}(r, 0)= \begin{cases}\sqrt{\frac{1}{2} \pi k r} J_{m}(k r) & \text { when } E>0 \text { and } k=\sqrt{E} \\ \mathrm{e}^{-\mathrm{i} m \pi / 2} \sqrt{\frac{1}{2} \pi \kappa r} J_{m}(\mathrm{i} \kappa r) & \text { when } E \leqslant 0 \text { and } \kappa=\sqrt{-E}\end{cases}$
where the factor in front of the radial function $R_{E m}(r)$ is not important. The solution $R_{E m}(r, 0)$ given in (11) is a real function. The logarithmic derivative at $r_{0}-$ for $E \leqslant 0$ is
$A_{m}(E, 0) \equiv\left\{\frac{1}{R_{E m}(r, 0)} \frac{\partial R_{E m}(r, 0)}{\partial r}\right\}_{r=r_{0}-}= \begin{cases}\frac{k J_{m}^{\prime}\left(k r_{0}\right)}{J_{m}\left(k r_{0}\right)}-\frac{1}{2 r_{0}} & \text { when } E>0 \\ \frac{\mathrm{i} \kappa J_{m}^{\prime}\left(\mathrm{i} \kappa r_{0}\right)}{J_{m}\left(\mathrm{i} \kappa r_{0}\right)}-\frac{1}{2 r_{0}} & \text { when } E \leqslant 0 .\end{cases}$
In the region $\left[r_{0}, \infty\right)$, we have $V(r)=U\left(r, r^{\prime}\right)=0$. For $E>0$, there are two oscillatory solutions to (6). Their combination can always satisfy the matching condition (7), so that there is a continuous spectrum for $E>0$ :

$$
\begin{align*}
R_{E m}(r, \lambda)= & \sqrt{\frac{1}{2} \pi k r}\left\{\cos \eta_{m}(k, \lambda) J_{m}(k r)-\sin \eta_{m}(k, \lambda) N_{m}(k r)\right\} \\
& \sim \cos \left(k r-\frac{1}{2} m \pi-\frac{1}{4} \pi+\eta_{m}(k, \lambda)\right) \quad \text { when } r \longrightarrow \infty \tag{13}
\end{align*}
$$

where $N_{m}(k r)$ is the Neumann function. The phase shift $\eta_{m}(k, \lambda)$ is determined by the matching condition (7):

$$
\begin{align*}
& \tan \eta_{m}(k, \lambda)=\frac{J_{m}\left(k r_{0}\right)}{N_{m}\left(k r_{0}\right)} \frac{A_{m}(E, \lambda)-k J_{m}^{\prime}\left(k r_{0}\right) / J_{m}\left(k r_{0}\right)-1 /\left(2 r_{0}\right)}{A_{m}(E, \lambda)-k N_{m}^{\prime}\left(k r_{0}\right) / N_{m}\left(k r_{0}\right)-1 /\left(2 r_{0}\right)}  \tag{14}\\
& \eta_{m}(k) \equiv \eta_{m}(k, 1) \tag{15}
\end{align*}
$$

where the primes denote the derivative of the Bessel, Neumann and later the Hankel function with respect to their argument. Although the radial equation (6) in the region $\left[r_{0}, \infty\right)$ is independent of $\lambda$, the solution $R_{E m}(r, \lambda)$ and the phase shift $\eta_{m}(k, \lambda)$ do depend on $\lambda$ through the matching condition (7).

The phase shift $\eta_{m}(k, \lambda)$ is determined from (14) up to a multiple of $\pi$ due to the period of the tangent function. Levinson determined the phase shift $\eta_{m}(k)$ with respect to the phase shift $\eta_{m}(\infty)$ at infinite momentum. For any finite potential, the phase shift $\eta_{m}(\infty)$ will not change and is always equal to zero. Therefore, Levinson's definition for the phase shift is equivalent to the convention that the phase shift $\eta_{m}(k, \lambda)$ is determined with respect to the phase shift $\eta_{m}(k, 0)$ for the free particle, where $\eta_{m}(k, 0)$ is defined to be zero:

$$
\begin{equation*}
\eta_{m}(k, 0)=0 \quad \text { where } \lambda=0 \tag{16}
\end{equation*}
$$

There is some ambiguity in $\eta_{m}(\infty)$ when a bound state with a positive energy occurs (see section 4). However, as far as the Levinson theorem is concerned, the latter convention is more convenient. We prefer to use this convention where the phase shift $\eta_{m}(k)$ is determined completely as $\lambda$ increases from zero to one. This is the reason why we introduce the parameter $\lambda$.

For $E \leqslant 0$ there is only one convergent solution at infinity

$$
\begin{equation*}
R_{E m}(r)=\mathrm{e}^{\mathrm{i}(m+1) \pi / 2} \sqrt{\frac{\pi \kappa r}{2}} H_{m}^{(1)}(\mathrm{i} \kappa r) \sim \mathrm{e}^{-\kappa r} \quad \text { when } r \longrightarrow \infty \tag{17}
\end{equation*}
$$

where $H_{m}^{(1)}(x)$ is the Hankel function of the first kind. Thus, the matching condition (7) is not always satisfied. When the matching condition (7) is satisfied, a bound state appears at this energy. It means that there is a discrete spectrum for $E \leqslant 0$. From equation (17) we have

$$
\begin{align*}
\left\{\frac{1}{R_{E m}(r, 0)}\right. & \left.\frac{\partial R_{E m}(r, 0)}{\partial r}\right\}_{r=r_{0}+}=\frac{\mathrm{i} \kappa H_{m}^{(1)}\left(\mathrm{i} \kappa r_{0}\right)^{\prime}}{H_{m}^{(1)}\left(\mathrm{i} \kappa r_{0}\right)}-\frac{1}{2 r_{0}} \\
& = \begin{cases}\left(-m+\frac{1}{2}\right) / r_{0} \equiv \rho_{m} & \text { when } E \longrightarrow 0 \\
-\kappa \sim-\infty & \text { when } E \longrightarrow-\infty\end{cases} \tag{18}
\end{align*}
$$

On the other hand, if $V(r)=U\left(r, r^{\prime}\right)=0$, from (12) we obtain
$A_{m}(E, 0) \equiv\left\{\frac{1}{R_{E m}(r, 0)} \frac{\partial R_{E m}(r, 0)}{\partial r}\right\}_{r=r_{0}-}$

$$
=\frac{\mathrm{i} \kappa J_{m}^{\prime}\left(\mathrm{i} \kappa r_{0}\right)}{J_{m}\left(\mathrm{i} \kappa r_{0}\right)}-\frac{1}{2 r_{0}}= \begin{cases}\left(m+\frac{1}{2}\right) / r_{0} & \text { when } E \longrightarrow 0  \tag{19}\\ \kappa \sim \infty & \text { when } E \longrightarrow-\infty\end{cases}
$$

It can easily be seen from (18) and (19) that as energy increases from $-\infty$ to 0 , there is no overlap between two variant ranges of the logarithmic derivatives either side of $r_{0}$ such that there is no bound state when $\lambda=0$ except for an $S$ wave where there is a half-bound state at $E=0$. The half-bound state will be discussed at the end of this section.

If $A_{m}(0, \lambda)$ decreases across the value $\rho_{m} \equiv\left(-m+\frac{1}{2}\right) / r_{0}$ as $\lambda$ increases, an overlap between two variant ranges of the logarithmic derivatives at two sides of $r_{0}$ appears. Since the logarithmic derivatives of the radial function at $r_{0}$ - decreases monotonically as the energy increases, and that at $r_{0}+$ increases monotonically, the overlap means that there must be one and only one energy where the matching condition (7) is satisfied, i.e. a scattering state changes to a bound state.

As $\lambda$ increases, a zero point in the zero-energy solution $R_{0 m}(r, \lambda)$ may occur through $r_{0}$. In this process $A_{m}(0, \lambda)$ may decrease to $-\infty$, jump to $+\infty$, and decreases again, or vice versa. It is not a singularity. If $A_{m}(0, \lambda)$ decreases through the jump at infinity, again across the value $\rho_{m}$, another bound state appears.

As $\lambda$ increases from zero to one, each time $A_{m}(0, \lambda)$ decreases across the value $\rho_{m}$, a new overlap between the variant ranges of two logarithmic derivatives appears such that a scattering state changes to a bound state. Conversely, each time $A_{m}(0, \lambda)$ increases across the value $\rho_{m}$, an overlap between those two variant ranges disappears such that a bound state changes back to a scattering state. The number of bound states $n_{m}$ is equal to the times that $A_{m}(0)$ decreases across the value $\rho_{m}$ as $\lambda$ change from zero to one, subtracted by the times that $A_{m}(0)$ increases across the value $\rho_{m}$. In wnat follows, we will show from (14) that this number is simply the phase shift $\eta_{m}(0)$ at zero momentum divided by $\pi$.

It can easily be seen from (14) that the phase shift $\eta_{m}(k, \lambda)$ increases monotonically as the logarithmic derivative $A_{m}(E)$ decreases:
$\left.\frac{\partial \eta_{m}(k, \lambda)}{\partial A_{m}(E, \lambda)}\right|_{k}=-\frac{8 r_{0} \cos ^{2} \eta_{m}(k)}{\pi\left(2 r_{0} A_{m}(E) N_{m}\left(k r_{0}\right)-2 k r_{0} N_{m}^{\prime}\left(k r_{0}\right)-N_{m}\left(k r_{0}\right)\right)^{2}} \leqslant 0$.

The phase shift $\eta_{m}(0, \lambda)$ is the limit of the phase shift $\eta_{m}(k, \lambda)$ as $k$ tends to zero. Therefore, what we are interested in is the phase shift $\eta_{m}(k, \lambda)$ at sufficiently small momentum $k$, $k \ll 1 / r_{0}$. For small momentum $k$, from (14) we obtain

$$
\begin{align*}
& \tan \eta_{m}(k, \lambda) \sim \\
& \begin{cases}\frac{-\pi\left(k r_{0}\right)^{2 m}}{2^{2 m} m!(m-1)!} \frac{A_{m}(0, \lambda)-\left(m+\frac{1}{2}\right) / r_{0}}{A_{m}(0, \lambda)-c^{2} k^{2}-\rho_{m}\left(1-\left(k r_{0}\right)^{2} /(m-1)(2 m-1)\right)} & \text { when } m \geqslant 2 \\
\frac{-\pi\left(k r_{0}\right)^{2}}{4} \frac{A_{m}(0, \lambda)-3 /\left(2 r_{0}\right)}{A_{m}(0, \lambda)-c^{2} k^{2}-\rho_{1}\left(1+2\left(k r_{0}\right)^{2} \ln \left(k r_{0}\right)\right)} & \text { when } m=1 \\
\frac{\pi}{2 \ln \left(k r_{0}\right)} \frac{A_{m}(0, \lambda)-c^{2} k^{2}-\rho_{0}\left(1-\left(k r_{0}\right)^{2}\right)}{A_{m}(0, \lambda)-c^{2} k^{2}-\rho_{0}\left(1+2 / \ln \left(k r_{0}\right)\right)} & \text { when } m=0\end{cases} \tag{21}
\end{align*}
$$

In addition to the leading terms, we include in (21) some next-to-leading terms, which are useful only for the critical case where the leading terms cancel each other out.

First, it can be seen from (21) that $\tan \eta_{m}(k, \lambda)$ tends to zero as $k$ goes to zero, i.e. $\eta_{m}(0, \lambda)$ is always equal to a multiple of $\pi$. In other words, if the phase shift $\eta_{m}(k, \lambda)$ for sufficiently small $k$ is expressed as a positive or negative acute angle plus $n \pi$, where $n$ is an integer, its limit $\eta_{m}(0, \lambda)$ is equal to $n \pi$. It means that $\eta_{m}(0, \lambda)$ changes discontinuously. By the way, in three dimensions, the phase shift at zero momentum of the $S$ wave may have an additional $\pi / 2$ when the half-bound state occurs.

Second, since the phase shift $\eta_{m}(k, \lambda)$ increases monotonically as the logarithmic derivative $A_{m}(E, \lambda)$ decreases, the phase shift at zero momentum $\eta_{m}(0, \lambda)$ will jump by $\pi$ if $\tan \eta_{m}(k, \lambda)$ at sufficiently small $k$ changes sign from positive to negative as $A_{m}(E, \lambda)$ decreases, and vice versa. When $\lambda$ changes from zero to one continuously, each time $A_{m}(0, \lambda)$ decreases from near and larger than the value $\rho_{m}$ to smaller than that value, the denominator in (21) changes sign from positive to negative and the remaining factor keeps positive, such that the phase shift at zero momentum $\eta_{m}(0, \lambda)$ jumps by $\pi$. Conversely, each time $A_{m}(0, \lambda)$ increases across the value $\rho_{m}$, the phase shift at zero momentum $\eta_{m}(0, \lambda)$ jumps by $-\pi$. Therefore, the phase shift $\eta_{m}(0) / \pi$ is just equal to the number of times $A_{m}(0, \lambda)$ decreases across the value $\rho_{m}$ as $\lambda$ increases from zero to one, minus the number of times $A_{m}(0, \lambda)$ increases across that value. Therefore, we have proved the Levinson theorem for the Schrödinger equation in two dimensions for non-critical cases:

$$
\begin{equation*}
\eta_{m}(0)=n_{m} \pi \tag{22a}
\end{equation*}
$$

Third, we should pay some attention to the case $m=0$. When $A_{0}(0, \lambda)$ decreases across the value $\rho_{0}=1 /\left(2 r_{0}\right)$, both the numerator and denominator in (21) change sign, but not spontaneously because the next-to-leading terms in the numerator and denominator in (21) are different. It is easy to see that the numerator changes sign first, and then the denominator changes sign, i.e. $\tan \eta_{0}(k, \lambda)$ at small $k$ changes first from negative to positive, then to negative again such that $\eta_{0}(0, \lambda)$ jumps by $\pi$. Similarly, when $A_{m}(0, \lambda)$ increases across the value $\rho_{0}, \eta_{0}(0, \lambda)$ jumps by $-\pi$.

When $\lambda=0$ and $m=0$, the numerator in (21) is equal to zero, the denominator is positive, and the phase shift $\eta_{0}(0)$ is defined to be zero. If $A_{0}(E)$ decreases when $\lambda$ increases from zero, the numerator becomes negative first, and then the denominator changes from positive to negative such that the phase shift $\eta_{0}(0)$ jumps by $\pi$ and a new bound state appears simultaneously.

Finally, we turn to discuss the critical cases where a half-bound state occurs. If the logarithmic derivative $A_{m}(0,1)$ is equal to the value $\rho_{m}$, the following solution with zero energy in the region $\left[r_{0}, \infty\right)$ will match this $A_{m}(0,1)$ at $r_{0}$ :

$$
R_{0 m}(r, 1)=r^{-m+1 / 2}
$$

It is a bound state when $m \geqslant 2$, but called a half-bound state when $m=1$ or 0 . A halfbound state is a zero-energy solution of the Schrödinger equation which is finite but which does not decay fast enough at infinity to be square integrable. We are going to discuss the critical case where $A_{m}(0, \lambda)$ decreases (or increases) and reaches, but does not cross, the value $\rho_{m}$ as $\lambda$ increases from a value somewhat smaller than one up to one. For definiteness, we discuss the case where $A_{m}(0, \lambda)$ decreases and reaches the value $\rho_{m}$ as $\lambda$ increases to one. In this case a new bound state with zero energy appears for $m \geqslant 2$, but does not appear for $m=1$ and 0 . We need to check whether or not the phase shift $\eta_{m}(0)$ increases an additional $\pi$.

It is evident that the denominator in (21) for $m \geqslant 2$ has changed sign from positive to negative as $A_{m}(0)$ decreases and reaches the value $\rho_{m}$, i.e. the phase shift $\eta_{m}(0)$ jumps by $\pi$ and simultaneously a new bound state of zero energy appears.

For $m=0$ the next-to-leading term of the denominator in (21) is positive and larger than the term $-c^{2} k^{2}$, such that the denominator does not change sign, i.e. the phase shift $\eta_{m}(0)$ does not jump. Simultaneously, no new bound state appears.

For $m=1$ the next-to-leading term of the denominator in (21) is negative such that the denominator does change sign, i.e. the phase shift $\eta_{m}(0)$ jumps by $\pi$ as $A_{m}(0)$ decreases and reaches the value $\rho_{1}$. However, in this case no new bound state appears simultaneously.

The discussion for the cases where $A_{m}(0)$ increases and reaches the value $\rho_{m}$ is similar. Therefore, Levinson's theorem ( $22 a$ ) holds for the critical cases except for $m=1$. In the latter case, Levinson's theorem for the Schrödinger equation with both local and non-local interactions in two dimensions becomes

$$
\begin{equation*}
\eta_{m}(0)=\left(n_{m}+1\right) \pi \quad \text { when } m=1 \quad \text { and a half-bound state occurs. } \tag{22b}
\end{equation*}
$$

As discussed above, it is found that the Levinson theorem holds without any modification for the case where a non-local potential is included.

## 4. Positive-energy bound states

It is well known that, in the case with only a local interaction, the wavefunction and its first derivative would never vanish at the same point except at the origin, so there is no positiveenergy bound state. However, in the case with a non-local interaction, Martin showed that the solution with an asymptotic form is not unique when the potential satisfies some conditions [4], i.e. there exists the positive-energy bound state with a vanishing asymptotic form. If a small perturbative potential is added such that the non-local potential satisfies the conditions, the positive-energy bound state will appear and the phase shift at this energy increases rapidly by almost $\pi$. This can be seen explicitly in the examples given by Martin [4] and Kermode [23].

It was pointed out by Kermode that the inverse tangent function is not single-valued and it is physically more satisfactory to include a jump of $\pi$ to the phase shift at the energy $E_{0}$, where a positive-energy bound state occurs. Martin and Chadan [4, 21] defined the phase shift to be continuous even at $E_{0}$ so that an additional $\pi$ will be included into $\delta(0)-\delta(\infty)$ for each positive-energy bound state. This is their reason for modifying Levinson's theorem
by the term $\sigma \pi$ where $\sigma$ denotes the number of the positive-energy bound states. However, according to the viewpoint of Kermode, no modification to the Levinson theorem is required.

However, the phase shift at zero energy in our convention does not change, no matter which viewpoint is used, i.e. no matter whether the phase shift jumps or not at the energy with a positive-energy bound state. Therefore, the Levinson theorem (22) holds for the cases where positive-energy bound states may occur.

## 5. The physically redundant state

The resonating group model of the scattering of nuclei, or other composite systems, derives an effective two-body interaction in which a non-local potential appears. There are some physically redundant states which describe Pauli-forbidden states for the compound system, and the physical two-body states must be orthogonal to these redundant states [24]. In the case of three dimensions, Saito [25], Okai et al [26], and Englefield and Shoukry [27] proposed a simple non-local term which guarantees the required orthogonality, and verified that it was a good representation of the interactions. If there is just one redundant state represented by the real normalized wavefunction $U(r)$, then the two-dimensional Saito equation is

$$
\begin{align*}
\frac{\mathrm{d}^{2}}{\mathrm{~d} r^{2}} R_{E m}(r) & +\left(E-V(r)-\frac{m^{2}-\frac{1}{4}}{r^{2}}\right) R_{E m}(r) \\
& =U(r) \int_{0}^{\infty} U(s)\left(\frac{\mathrm{d}^{2}}{\mathrm{~d} s^{2}}-V(s)-\frac{m^{2}-\frac{1}{4}}{s^{2}}\right) R_{E m}(s) \mathrm{d} s \tag{23}
\end{align*}
$$

$\int_{0}^{\infty} U^{2}(s) \mathrm{d} s=1$
and

$$
\begin{equation*}
E \int_{0}^{\infty} U(r) R_{E m}(r) \mathrm{d} r=0 \tag{24}
\end{equation*}
$$

The solution of (23) satisfies the orthogonality constraint except for that of zero energy. Saito's non-local potential is separable.

If the Schrödinger equation with only a local potential $V(r)$ has a bound state with a negative $-\mathcal{E}<0$, the corresponding wavefunction is denoted by $\psi(r)$ :

$$
\begin{align*}
& \frac{\mathrm{d}^{2}}{\mathrm{~d} r^{2}} \psi(r)-\left(V(r)+\frac{m^{2}-\frac{1}{4}}{r^{2}}\right) \psi(r)=\mathcal{E} \psi(r)  \tag{25}\\
& \int_{0}^{\infty} \psi(r)^{2} \mathrm{~d} r=1
\end{align*}
$$

It is obvious that $U(r)=\psi(r)$ satisfies (23) with zero energy. Therefore, it is the so-called physically redundant state. As far as equation (23) is concerned, the redundant state is one of the bound states with zero energy.

## Acknowledgments

This work was supported by the National Natural Science Foundation of China and Grant No LWTZ-1298 of the Chinese Academy of Sciences.

## References

[1] Levinson N 1949 K. Danske Vidensk. Selsk., Mat.-Fys. Medd. 25 (9)
[2] Newton R G 1960 J. Math. Phys. 1319
Newton R G 1977 J. Math. Phys. 181348
Newton R G 1977 J. Math. Phys. 181582
Newton R G 1982 Scattering Theory of Waves and Particles 2nd edn (Berlin: Springer) and references therein
[3] Jauch J M 1957 Helv. Phys. Acta 30143
[4] Martin A 1958 Nuovo Cimento 7607
[5] Ma Z Q and Dai A Y 1988 J. Phys. A: Math. Gen. 212085
[6] Ni G J 1979 Phys. Energ. Fort. Phys. Nucl. 3432
Ma Z Q and Ni G J 1985 Phys. Rev. D 311482
Ma Z Q 1995 J. Math. Phys. 261995
[7] Ma Z Q 1985 Phys. Rev. D 322203
Ma Z Q 1985 Phys. Rev. D 322213
[8] Iwinski Z R, Rosenberg L and Spruch L 1985 Phys. Rev. 311229
[9] Poliatzky N 1993 Phys. Rev. Lett. 702507
Newton R G 1994 Helv. Phys. Acta 6720
Ma Z Q 1996 Phys. Rev. Lett. 763654
[10] Dong Shi-Hai, Hou Xi-Wen and Ma Zhong-Qi 1998 Levinson's theorem for the Schrödinger equation in two dimensions Phys. Rev. A, accepted
[11] Dong Shi-Hai, Hou Xi-Wen and Ma Zhong-Qi 1998 The relativistic Levinson's theorem in two dimensions Phys. Rev. A, accepted
[12] Portnoi M E and Galbraith I 1997 Solid State Commun. 103325
[13] Bollé D, Gesztesy F, Danneels C and Wilk S F J 1986 Phys. Rev. Lett. 56900
[14] Iwinski Z R, Rosenberg L and Spruch L 1986 Phys. Rev. A 33946 Rosenberg L and Spruch L 1996 Phys. Rev. A 544985
[15] Blankenbecler R and Boyanovsky D 1986 Physica 18D 367
[16] Niemi A J and Semenoff G W 1985 Phys. Rev. D 32471
[17] Vidal F and Letourneaux J 1992 Phys. Rev. C 45418
[18] Kiers K A and van Dijk W 1996 J. Math. Phys. 376033
[19] Debianchi M S 1994 J. Math. Phys. 352719
[20] Martin P A and Debianchi M S 1996 Europhys. Lett. 34639
[21] Chadan Kh 1958 Nuovo Cimento 10892
[22] Yang C N 1982 Monopoles in Quantum Field Theory (Proc. Monopole Meeting, Trieste, Italy, 1981) ed N S Craigie, P Goddard and W Nahm (Singapore: World Scientific) p 237
[23] Kermode M W 1976 J. Phys. A: Math. Gen. 9 L47
[24] Tamagaki R 1968 Prog. Theor. Phys. Suppl. 242 242-58
[25] Saito S 1968 Prog. Theor. Phys. 40893
Saito S 1969 Prog. Theor. Phys. 41705
[26] Okai S, Saito S and Tamagaki R 1972 Prog. Theor. Phys. 47484
[27] Englefield M J and Shoukry H S M 1974 Prog. Theor. Phys. 521554


[^0]:    || E-mail: dongsh@bepc4.ihep.ac.cn

